



1
00:00:00,790 --> 00:00:07,320

[Music]

2
00:00:12,379 --> 00:00:09,080

[Applause]

3
00:00:15,869 --> 00:00:12,389

I'm gonna tell you about these detection

4
00:00:17,790 --> 00:00:15,879

that we reported in the paper about year

5
00:00:20,250 --> 00:00:17,800

and a half ago of methyl chloride in

6
00:00:22,140 --> 00:00:20,260

particular and the co materia regions

7
00:00:24,450 --> 00:00:22,150

and also about the ongoing lab work that

8
00:00:26,310 --> 00:00:24,460

we're doing at JPL to support this

9
00:00:29,069 --> 00:00:26,320

detection and maybe guide future

10
00:00:31,319 --> 00:00:29,079

observation to observe more of these

11
00:00:34,319 --> 00:00:31,329

pieces so first of all why methyl

12
00:00:36,660 --> 00:00:34,329

halides well these pieces are abundant

13
00:00:39,150 --> 00:00:36,670

in the atmosphere of the earth so a

14

00:00:40,860 --> 00:00:39,160

couple of decades ago they were mainly

15

00:00:43,110 --> 00:00:40,870

found because of industrial processes

16

00:00:44,670 --> 00:00:43,120

but these days they're actually the the

17

00:00:47,569 --> 00:00:44,680

main production source is actually for

18

00:00:51,209 --> 00:00:47,579

biological processes so for example

19

00:00:55,650 --> 00:00:51,219

metabolizing allergens into the oceans

20

00:00:58,669 --> 00:00:55,660

by allergies and they will reject our

21

00:01:01,709 --> 00:00:58,679

gano allergens into the atmosphere and

22

00:01:05,490 --> 00:01:01,719

due to this connection between biology

23

00:01:08,610 --> 00:01:05,500

and for these pieces they were proposed

24

00:01:12,480 --> 00:01:08,620

as potential biomarkers if observed in

25

00:01:14,310 --> 00:01:12,490

the atmosphere of exoplanets this was

26

00:01:15,420 --> 00:01:14,320

further supported by the fact that in

27

00:01:18,450 --> 00:01:15,430

the interstellar medium

28

00:01:21,840 --> 00:01:18,460

well we actually detect allergens pieces

29

00:01:24,530 --> 00:01:21,850

but very small ones so HCl HCl plus and

30

00:01:26,880 --> 00:01:24,540

so on and most of the halogens

31

00:01:28,800 --> 00:01:26,890

especially for chlorine here is actually

32

00:01:31,410 --> 00:01:28,810

found in the diffuse interstellar medium

33

00:01:34,170 --> 00:01:31,420

and noting two molecular clouds were

34

00:01:35,550 --> 00:01:34,180

star formations happening so the

35

00:01:37,830 --> 00:01:35,560

question we have is there is there an

36

00:01:39,810 --> 00:01:37,840

actual disconnect between the allergen

37

00:01:43,740 --> 00:01:39,820

chemistry and the organic chemistry that

38

00:01:47,190 --> 00:01:43,750

we see in star forming and planet forming

39

00:01:50,550 --> 00:01:47,200

environments so to figure this out we

40

00:01:53,100 --> 00:01:50,560

search for them into purchased of our

41

00:01:55,950 --> 00:01:53,110

environments and specifically within the

42

00:01:58,859 --> 00:01:55,960

framework of the pills program but is an

43

00:02:01,230 --> 00:01:58,869

Alma program so Alma looked at the

44

00:02:03,270 --> 00:02:01,240

rotation lines towards a low mass

45

00:02:06,990 --> 00:02:03,280

produced are here for this pills program

46

00:02:09,389 --> 00:02:07,000

led by yes or Genson so this producer RS

47

00:02:11,640 --> 00:02:09,399

16 to 9 3 24 22 is actually thought to

48

00:02:14,360 --> 00:02:11,650

be an analogue of water sun looked like

49

00:02:17,600 --> 00:02:14,370

when it was forming so we looked at

50

00:02:20,420 --> 00:02:17,610

rotational lines nearby one of

51
00:02:23,000 --> 00:02:20,430
sources source be around 16 to 9 free

52
00:02:25,400 --> 00:02:23,010
and within this large program we

53
00:02:27,890 --> 00:02:25,410
actually covered a large frequency range

54
00:02:32,180 --> 00:02:27,900
at a very nice pectoral resolution that

55
00:02:35,120 --> 00:02:32,190
allowed us to to look at this lines and

56
00:02:37,310 --> 00:02:35,130
also at a very high angular resolution

57
00:02:39,080 --> 00:02:37,320
point five our second which for this

58
00:02:40,490 --> 00:02:39,090
sources correspond to about sixty you

59
00:02:43,699 --> 00:02:40,500
it's actually very small for pretty

60
00:02:46,040 --> 00:02:43,709
stuff so this is what the data look like

61
00:02:47,780 --> 00:02:46,050
it's lots and lots of lines it's

62
00:02:50,870 --> 00:02:47,790
actually a spectrum that is taking at

63
00:02:52,400 --> 00:02:50,880

one special like expected at one

64

00:02:54,620 --> 00:02:52,410

position towards the star we actually

65

00:02:56,590 --> 00:02:54,630

have a whole map of these and all of

66

00:03:00,650 --> 00:02:56,600

these are rotational lines from

67

00:03:02,360 --> 00:03:00,660

molecules and myths of them are like

68

00:03:03,680 --> 00:03:02,370

molecules that we know that we have been

69

00:03:06,860 --> 00:03:03,690

observed that have been observed before

70

00:03:09,440 --> 00:03:06,870

like methanol dimethyl ether and so on

71

00:03:10,790 --> 00:03:09,450

deteriorated version of these but some

72

00:03:12,320 --> 00:03:10,800

of these lines are actually due to

73

00:03:15,080 --> 00:03:12,330

molecules that haven't been identified

74

00:03:18,140 --> 00:03:15,090

before and this was the case for methyl

75

00:03:20,300 --> 00:03:18,150

chloride so this is what we found this

76

00:03:23,960 --> 00:03:20,310

is the detection that we got so for the

77

00:03:26,570 --> 00:03:23,970

35 chlorine and the 37 chlorine as you

78

00:03:30,560 --> 00:03:26,580

can see the lines are like very intense

79

00:03:33,350 --> 00:03:30,570

we can actually detect them up to k

80

00:03:38,090 --> 00:03:33,360

equal four into is 13 to 12 transition

81

00:03:40,280 --> 00:03:38,100

for this symmetric top molecule so nice

82

00:03:42,110 --> 00:03:40,290

we can fit them we can derive a benign

83

00:03:44,770 --> 00:03:42,120

we can direct column density for these

84

00:03:48,710 --> 00:03:44,780

two species and by doing the ratio of 35

85

00:03:50,660 --> 00:03:48,720

chlorine over 37 we actually find 2.1 in

86

00:03:52,370 --> 00:03:50,670

this star forming region so for those of

87

00:03:54,470 --> 00:03:52,380

you who are familiar with the value of

88

00:03:56,060 --> 00:03:54,480

isotopic ratio for chlorine and the

89

00:03:57,590 --> 00:03:56,070

solar system it's going to be a bit of a

90

00:04:00,410 --> 00:03:57,600

surprise in the solar system we expect

91

00:04:03,440 --> 00:04:00,420

the value at 3.1 but it's actually not

92

00:04:05,570 --> 00:04:03,450

surprising to find 2.1 because detection

93

00:04:08,330 --> 00:04:05,580

previous detection of HC else of it

94

00:04:11,449 --> 00:04:08,340

found this ratio to be 2.1 and it's

95

00:04:14,509 --> 00:04:11,459

actually now accepted that this ratio

96

00:04:19,219 --> 00:04:14,519

changes in different molecular clouds

97

00:04:22,280 --> 00:04:19,229

nearby our solar system so this is what

98

00:04:25,490 --> 00:04:22,290

also combining these two abundances and

99

00:04:27,150 --> 00:04:25,500

comparing them to that of methanol one

100

00:04:29,550 --> 00:04:27,160

of very big

101
00:04:32,580 --> 00:04:29,560
this organic molecule in this region we

102
00:04:34,890 --> 00:04:32,590
found a ratio about 10 to the minus 4 so

103
00:04:36,630 --> 00:04:34,900
it's Medicare ID is not in abundance

104
00:04:38,850 --> 00:04:36,640
pieces in these environments but it's

105
00:04:40,920 --> 00:04:38,860
not negligible either and we see it here

106
00:04:44,160 --> 00:04:40,930
I'm showing you here the emission of

107
00:04:47,340 --> 00:04:44,170
methyl chloride around this protostars

108
00:04:48,540 --> 00:04:47,350
eschews be here and I was actually the

109
00:04:53,370 --> 00:04:48,550
spectrum that I show you here was

110
00:04:55,140 --> 00:04:53,380
extracted at this position to here so so

111
00:04:58,230 --> 00:04:55,150
we detected this methyl chloride around

112
00:05:00,180 --> 00:04:58,240
illumise protostar and some of the

113
00:05:02,490 --> 00:05:00,190

members of the pills team we're actually

114

00:05:05,670 --> 00:05:02,500

also members of the Rosina team that

115

00:05:09,750 --> 00:05:05,680

looked at the outgassing a comet 67p and

116

00:05:12,120 --> 00:05:09,760

so we like them to look in the spectra

117

00:05:14,130 --> 00:05:12,130

of Rosina and especially the DFM s to

118

00:05:16,080 --> 00:05:14,140

see if a that we would find this methyl

119

00:05:19,560 --> 00:05:16,090

tried and it was actually there have a

120

00:05:22,470 --> 00:05:19,570

small shoulder here around my 50s so

121

00:05:24,170 --> 00:05:22,480

that the the resolution of this mass

122

00:05:27,990 --> 00:05:24,180

spectrometer is such that you can

123

00:05:31,410 --> 00:05:28,000

disentangle different isomer at one

124

00:05:33,510 --> 00:05:31,420

specific amu we can disentangle for

125

00:05:36,330 --> 00:05:33,520

different atomic composition and that's

126

00:05:39,000 --> 00:05:36,340

how we could like get this methyl

127

00:05:41,640 --> 00:05:39,010

chloride detection we see the 35

128

00:05:44,700 --> 00:05:41,650

chlorine we don't see the 37 but it's

129

00:05:48,780 --> 00:05:44,710

not surprising if you accept this free

130

00:05:51,240 --> 00:05:48,790

point one isotopic ratio for chlorine in

131

00:05:53,070 --> 00:05:51,250

the solar system what is interesting is

132

00:05:56,370 --> 00:05:53,080

to look at the abundance at which we

133

00:05:59,190 --> 00:05:56,380

observed ethyl chloride versus methanol

134

00:06:01,050 --> 00:05:59,200

around 67 P and we found 10 to the minus

135

00:06:04,440 --> 00:06:01,060

4 which is similar to what we've seen

136

00:06:06,980 --> 00:06:04,450

for the protostar and this supports the

137

00:06:09,360 --> 00:06:06,990

possible scenario of an inheritance of

138

00:06:11,790 --> 00:06:09,370

chemistry form in the pretty little

139

00:06:13,830 --> 00:06:11,800

regions that would be inherited down to

140

00:06:16,020 --> 00:06:13,840

come at from ingredients or at least a

141

00:06:19,530 --> 00:06:16,030

scenario what the chemistry is similar

142

00:06:21,210 --> 00:06:19,540

in these two environments so the

143

00:06:23,670 --> 00:06:21,220

question is how do we form methyl

144

00:06:25,770 --> 00:06:23,680

chlorine how does it got here and while

145

00:06:27,270 --> 00:06:25,780

we don't know yet but other chemical

146

00:06:30,420 --> 00:06:27,280

models actually took a crack headed

147

00:06:33,090 --> 00:06:30,430

after this detection was published and

148

00:06:35,279 --> 00:06:33,100

they actually found that you could form

149

00:06:37,529 --> 00:06:35,289

a methyl chloride efficiently by

150

00:06:39,980 --> 00:06:37,539

chlorination or at regeneration of

151

00:06:43,280 --> 00:06:39,990

carbon of methyl onto the gray

152

00:06:47,030 --> 00:06:43,290

surface that you find in this in the

153

00:06:50,900 --> 00:06:47,040

purchaser or envelope so you form CH_3Cl

154

00:06:53,480 --> 00:06:50,910

and then the grains in the outer part of

155

00:06:56,420 --> 00:06:53,490

the particular envelope flow towards the

156

00:06:59,120 --> 00:06:56,430

central object so when it's forming and

157

00:07:01,160 --> 00:06:59,130

so the grains warm up and Myka's that

158

00:07:03,500 --> 00:07:01,170

you would form into the the ice surface

159

00:07:06,200 --> 00:07:03,510

will dissolve in the gas phase and be

160

00:07:09,620 --> 00:07:06,210

able that's how we can detect them with

161

00:07:11,480 --> 00:07:09,630

radio interferometers of course this is

162

00:07:13,430 --> 00:07:11,490

the theory this is what the models tells

163

00:07:15,410 --> 00:07:13,440

us but we need to perform laboratory

164

00:07:17,810 --> 00:07:15,420

experiments to figure out what is the

165

00:07:19,970 --> 00:07:17,820

actual formation mechanism but before

166

00:07:23,510 --> 00:07:19,980

doing that what we also need to know is

167

00:07:25,220 --> 00:07:23,520

how to quantify methyl chloride in this

168

00:07:28,100 --> 00:07:25,230

in the solid state if we want to test

169

00:07:29,720 --> 00:07:28,110

this formation scenario and for this so

170

00:07:32,030 --> 00:07:29,730

we need to do laboratory study and

171

00:07:34,940 --> 00:07:32,040

specifically looking at a spectroscopic

172

00:07:37,760 --> 00:07:34,950

studies because spectroscopy is the main

173

00:07:40,430 --> 00:07:37,770

tool that we have to look at molecule

174

00:07:42,380 --> 00:07:40,440

formation in the solid state so infrared

175

00:07:43,790 --> 00:07:42,390

spectroscopy in the mid are so that's

176

00:07:46,550 --> 00:07:43,800

what we're currently doing so it's

177

00:07:48,410 --> 00:07:46,560

ongoing work at JPL

178

00:07:51,530 --> 00:07:48,420

that we're doing cycle looking in the

179

00:07:53,750 --> 00:07:51,540

spectroscopy of these methyl halides so

180

00:07:55,430 --> 00:07:53,760

in pure ice films in this case and what

181

00:07:58,910 --> 00:07:55,440

we would like to know is to derive band

182

00:08:02,270 --> 00:07:58,920

strength for the different vibration of

183

00:08:03,710 --> 00:08:02,280

this pure of its molecules in the solid

184

00:08:05,960 --> 00:08:03,720

States so here I'm showing you the

185

00:08:09,830 --> 00:08:05,970

spectra that we get by depositing so

186

00:08:12,220 --> 00:08:09,840

molecule so CH_3FC or a CLC Revere or C

187

00:08:15,170 --> 00:08:12,230

a tree I on two cold substrates

188

00:08:16,370 --> 00:08:15,180

interrogating the spectroscopy with

189

00:08:19,430 --> 00:08:16,380

Fourier transform infrared spectroscopy

190

00:08:24,020 --> 00:08:19,440

and as you can see well we can attribute

191

00:08:25,820 --> 00:08:24,030

the emission the vibrations so for the

192

00:08:26,900 --> 00:08:25,830

CH stretching and the CH free

193

00:08:29,000 --> 00:08:26,910

deformation here

194

00:08:31,250 --> 00:08:29,010

there's not a huge change between the

195

00:08:33,260 --> 00:08:31,260

different types of allergens but of

196

00:08:37,850 --> 00:08:33,270

course once you go towards here tree

197

00:08:40,790 --> 00:08:37,860

rocks or CAC X so C F or C CL stretch

198

00:08:42,470 --> 00:08:40,800

very clear like changes due to the

199

00:08:46,160 --> 00:08:42,480

presence of the allergen and the nature

200

00:08:47,450 --> 00:08:46,170

of the allergen to get the bends ranked

201

00:08:50,060 --> 00:08:47,460

what we're doing is that we're

202

00:08:51,770 --> 00:08:50,070

integrating the optical depth under

203

00:08:53,460 --> 00:08:51,780

these bands divided by the column

204

00:08:56,370 --> 00:08:53,470

density of ice between

205

00:08:58,170 --> 00:08:56,380

and this is what we report here I warmed

206

00:09:00,690 --> 00:08:58,180

up the ice at different temperatures and

207

00:09:03,510 --> 00:09:00,700

as you can see the band strength for a

208

00:09:05,760 --> 00:09:03,520

difference see extra edge and CHP rock

209

00:09:09,180 --> 00:09:05,770

doesn't really change with temperature

210

00:09:11,040 --> 00:09:09,190

I'm inversa decrees here for chri high

211

00:09:13,890 --> 00:09:11,050

temperature most likely due to the onset

212

00:09:16,020 --> 00:09:13,900

of this origin but what actually changed

213

00:09:18,600 --> 00:09:16,030

is the shape of the features that you

214

00:09:20,880 --> 00:09:18,610

get so this is what we get when we

215

00:09:22,710 --> 00:09:20,890

deposit this films at low temperature

216

00:09:25,500 --> 00:09:22,720

and when we increase temperature we have

217

00:09:27,680 --> 00:09:25,510

like like very different shape my

218

00:09:31,170 --> 00:09:27,690

Seikaly due to the crystallization of

219

00:09:33,090 --> 00:09:31,180

these films in water ice we also observe

220

00:09:34,740 --> 00:09:33,100

the change of temperature here that is

221

00:09:38,820 --> 00:09:34,750

most likely due to the diffusion of

222

00:09:41,070 --> 00:09:38,830

molecules so here so water to see a tree

223

00:09:42,990 --> 00:09:41,080

I attend to one mixture at high

224

00:09:44,880 --> 00:09:43,000

temperature you can have diffusion these

225

00:09:46,410 --> 00:09:44,890

molecules likely segregates and that's

226

00:09:50,430 --> 00:09:46,420

why we can observe signature of

227

00:09:51,090 --> 00:09:50,440

crystallization so we get this been

228

00:09:53,940 --> 00:09:51,100

strength

229

00:09:55,800 --> 00:09:53,950

another characteristic and other

230

00:09:57,510 --> 00:09:55,810

behavior of this icing that we are

231

00:10:01,140 --> 00:09:57,520

highly interested in would be to look at

232

00:10:02,730 --> 00:10:01,150

the thermal desorption of his ices and

233

00:10:06,480 --> 00:10:02,740

that's what we've been doing this time

234

00:10:08,610 --> 00:10:06,490

by depositing ice film onto not a window

235

00:10:12,440 --> 00:10:08,620

band to a currently called

236

00:10:14,820 --> 00:10:12,450

micro balance so we deposited them

237

00:10:17,100 --> 00:10:14,830

different ice films of different

238

00:10:20,010 --> 00:10:17,110

thicknesses and what we did with micro

239

00:10:21,480 --> 00:10:20,020

balance is to follow the ice loss versus

240

00:10:24,180 --> 00:10:21,490

temperature and looking at the

241

00:10:25,980 --> 00:10:24,190

derivative here so ice loss rate versus

242

00:10:27,750 --> 00:10:25,990

temperature which you can see is that

243

00:10:29,820 --> 00:10:27,760

for the different thicknesses well the

244

00:10:32,580 --> 00:10:29,830

rate does not change this is typical of

245

00:10:36,770 --> 00:10:32,590

zero for the desorption kinetics you can

246

00:10:39,570 --> 00:10:36,780

fit this rates by music by this equation

247

00:10:40,860 --> 00:10:39,580

extracting here the doozers and energy

248

00:10:43,560 --> 00:10:40,870

said adsorption that you need to

249

00:10:46,860 --> 00:10:43,570

overcome to have desorption occurring

250

00:10:48,060 --> 00:10:46,870

these are the energy that I gets as you

251

00:10:50,700 --> 00:10:48,070

can see if there's not a clear

252

00:10:52,680 --> 00:10:50,710

correlation with the mass here Tracy L

253

00:10:54,540 --> 00:10:52,690

is a bit out of the way here and this

254

00:10:56,370 --> 00:10:54,550

miss actually because not just the mass

255

00:10:59,670 --> 00:10:56,380

is involved into the desorption of this

256

00:11:02,040 --> 00:10:59,680

species but also other parameters such

257

00:11:04,530 --> 00:11:02,050

as the polarity or even the size of

258

00:11:05,320 --> 00:11:04,540

these molecules CH3Cl here is really

259

00:11:07,120 --> 00:11:05,330

standing out

260

00:11:09,040 --> 00:11:07,130

could he explain this like the dipole

261

00:11:10,870 --> 00:11:09,050

moment here and could explain why it's

262

00:11:15,940 --> 00:11:10,880

stickier and it takes more energy to

263

00:11:18,130 --> 00:11:15,950

dissolve so pushing further towards

264

00:11:20,110 --> 00:11:18,140

Astrophysical implication we can use

265

00:11:23,860 --> 00:11:20,120

this sublimation energy that we derived

266

00:11:26,620 --> 00:11:23,870

in the laboratory to like measure well

267

00:11:28,930 --> 00:11:26,630

predicts and derive where the desorption

268

00:11:30,819 --> 00:11:28,940

France of these species are going to be

269

00:11:33,670 --> 00:11:30,829

in around protostars

270

00:11:35,380 --> 00:11:33,680

so like derive where the snow lines are

271

00:11:37,569 --> 00:11:35,390

going to occur and that's what I've been

272

00:11:39,100 --> 00:11:37,579

doing here with quick calculation so the

273

00:11:41,019 --> 00:11:39,110

definition of the snow line of the

274

00:11:44,590 --> 00:11:41,029

desorption front is where the flux of

275

00:11:48,430 --> 00:11:44,600

adsorption equals the flux of adsorption

276

00:11:50,620 --> 00:11:48,440

equals the flux of desorption by like

277

00:11:53,110 --> 00:11:50,630

equating the impingement rate of the

278

00:11:55,269 --> 00:11:53,120

item of the molecules equals to

279

00:11:57,519 --> 00:11:55,279

desorption rates you can derive this

280

00:12:00,400 --> 00:11:57,529

temperature here which for a species

281

00:12:02,650 --> 00:12:00,410

using typical abundances gives you the

282

00:12:05,319 --> 00:12:02,660

following temperature around the

283

00:12:09,699 --> 00:12:05,329

protostar and by like using this

284

00:12:11,500 --> 00:12:09,709

temperature and a temperature profile

285

00:12:14,410 --> 00:12:11,510

around Purdue star so here it's actually

286

00:12:15,670 --> 00:12:14,420

the profile for sixty to ninety three we

287

00:12:17,530 --> 00:12:15,680

are observed methyl chloride you can

288

00:12:19,180 --> 00:12:17,540

actually predict that methyl chloride

289

00:12:22,090 --> 00:12:19,190

pure methyl chloride will dissolve

290

00:12:24,220 --> 00:12:22,100

around 130 EU from the central object

291

00:12:28,210 --> 00:12:24,230

it's going to be further out for methyl

292

00:12:30,819 --> 00:12:28,220

fluoride and closer for the other methyl

293

00:12:34,329 --> 00:12:30,829

halide which is interesting to see also

294

00:12:34,720 --> 00:12:34,339

is that another 140 you value better

295

00:12:36,610 --> 00:12:34,730

again

296

00:12:38,470 --> 00:12:36,620

it's actually quite close to where we

297

00:12:40,360 --> 00:12:38,480

detect this methyl chloride here with

298

00:12:42,040 --> 00:12:40,370

our mother beam that I show you here is

299

00:12:44,199 --> 00:12:42,050

about 68 you

300

00:12:46,540 --> 00:12:44,209

and it's actually corresponds to here

301
00:12:48,519 --> 00:12:46,550
where the two beams where here we have

302
00:12:50,680 --> 00:12:48,529
this like onset of this urgent when you

303
00:12:53,860 --> 00:12:50,690
start to see methyl chloride popping up

304
00:12:58,269 --> 00:12:53,870
into the gas phase all right so that

305
00:13:01,210 --> 00:12:58,279
leaves me to my conclusions so I showed

306
00:13:02,980 --> 00:13:01,220
you detections of methyl chloride around

307
00:13:05,829 --> 00:13:02,990
pretty star around pretty star and

308
00:13:09,370 --> 00:13:05,839
around a comet's really showing that

309
00:13:11,439 --> 00:13:09,380
there's a link between the organic and

310
00:13:13,900 --> 00:13:11,449
the allergen chemistry in industrial

311
00:13:16,980 --> 00:13:13,910
environments these molecules are not

312
00:13:19,740 --> 00:13:16,990
direct biomarkers if observed in the

313
00:13:21,750 --> 00:13:19,750

in the atmosphere of exempt annette's

314

00:13:23,220 --> 00:13:21,760

however we need more detection to figure

315

00:13:25,710 --> 00:13:23,230

out what is the standard like the

316

00:13:28,740 --> 00:13:25,720

regular like typical amounts of his

317

00:13:30,240 --> 00:13:28,750

compounds in interest our environment to

318

00:13:36,389 --> 00:13:30,250

see if an exoplanet she would have

319

00:13:38,579 --> 00:13:36,399

differences in in in in abundance for

320

00:13:41,310 --> 00:13:38,589

the spectroscopic experiments we serve

321

00:13:44,550 --> 00:13:41,320

at the CX and the CH works were highly

322

00:13:46,620 --> 00:13:44,560

depending on the type of methyl halides

323

00:13:48,269 --> 00:13:46,630

that you have and the ur like depends

324

00:13:51,180 --> 00:13:48,279

their shape depends on the temperature

325

00:13:52,769 --> 00:13:51,190

of the ice film so they can be used to

326

00:13:55,040 --> 00:13:52,779

probe the ice environment we also derive

327

00:13:58,320 --> 00:13:55,050

Ben's rain that now can be employed to

328

00:14:00,720 --> 00:13:58,330

quantify a cone the formation of these

329

00:14:03,120 --> 00:14:00,730

pieces in ice films the desorption

330

00:14:04,949 --> 00:14:03,130

energy showed us that they're not

331

00:14:06,870 --> 00:14:04,959

circular it is to the mass of these

332

00:14:10,019 --> 00:14:06,880

pieces that other parameters are

333

00:14:13,050 --> 00:14:10,029

involved when predicting when measuring

334

00:14:15,240 --> 00:14:13,060

disturbs in energy of molecules you need

335

00:14:18,660 --> 00:14:15,250

to look at effects like the size of

336

00:14:20,160 --> 00:14:18,670

polarity and the calculation of where

337

00:14:21,420 --> 00:14:20,170

the sublimation front would be showed us

338

00:14:23,579 --> 00:14:21,430

that it's in agreement with what has

339

00:14:25,949 --> 00:14:23,589

been observed by Alma but the search for

340

00:14:28,110 --> 00:14:25,959

CH ret which didn't detect yet is going

341

00:14:30,180 --> 00:14:28,120

to be much more complicated since it is

342

00:14:32,360 --> 00:14:30,190

or further out and the Colombian cities

343

00:14:35,880 --> 00:14:32,370

are expecting dust to be lower than